

MOMENT EQUATIONS OF A DISTRIBUTION OF PARTICLES
PERFORMING HARMONIC OSCILLATIONS

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The moment equations have been studied by F. J. Sacherer (CERN/SI/Int.DL/70-12, 18.11.1970) up to the second moment. Here we extend it to higher moments.

Let the one-dimensional no space-charge single-particle equations be

$$\begin{cases} x' = p \\ p' = -kx \end{cases} \quad (\text{prime means } \frac{d}{dt}) \quad (1)$$

and the distribution function be $\psi(x, p, t)$ which obeys the continuity equation

$$\frac{\partial \psi}{\partial t} + \frac{\partial}{\partial x}(x' \psi) + \frac{\partial}{\partial p}(p' \psi) = 0 \quad (2)$$

For now we shall consider k to be time independent.

MOMENT EQUATIONS

The equations for the various moments of the distribution can be derived simply as follows:

First Moment

$$\left\{ \begin{aligned} (\bar{x})' &= \frac{d}{dt} \int x \psi(x, p, t) dx dp \\ &= \int x \frac{\partial \psi}{\partial t} dx dp \\ &= - \int x \left[\frac{\partial}{\partial x}(x' \psi) + \frac{\partial}{\partial p}(p' \psi) \right] dx dp \end{aligned} \right. \quad (3)$$



$$\left\{ \begin{array}{l} = \int x' \psi \, dx \, dp = \bar{x}' = \bar{p} \\ (\bar{p})' = \bar{p}' = -k\bar{x} \end{array} \right.$$

where we have performed a partial integration and made use of the assumption that $\psi \rightarrow 0$ as $|x| \rightarrow \infty$ or $|p| \rightarrow \infty$. We can write these two equations as

$$X_1' = K_1 X_1 \quad (4)$$

where

$$X_1 \equiv \begin{pmatrix} \bar{x} \\ \bar{p} \end{pmatrix} \quad K_1 \equiv \begin{pmatrix} 0 & 1 \\ -k & 0 \end{pmatrix}$$

Second Moment

$$\left\{ \begin{array}{l} (\bar{x}^2)' = 2 \bar{x} \bar{x}' = 2 \bar{x} \bar{p} \\ (\bar{x} \bar{p})' = \bar{x} \bar{p}' + \bar{x}' \bar{p} = -k \bar{x}^2 + \bar{p}^2 \\ (\bar{p}^2)' = 2 \bar{p} \bar{p}' = -2k \bar{x} \bar{p} \end{array} \right. \quad (5)$$

Or we can write

$$X_2' = K_2 X_2 \quad (6)$$

where

$$X_2 \equiv \begin{pmatrix} \bar{x}^2 \\ \bar{x} \bar{p} \\ \bar{p}^2 \end{pmatrix} \quad K_2 \equiv \begin{pmatrix} 0 & 2 & 0 \\ -k & 0 & 1 \\ 0 & -2k & 0 \end{pmatrix}$$

Similarly, we get the equations up to, say, the fifth moment. They are summarized as

$$X_n' = K_n X_n \quad (7)$$

where

$$X_1 = \begin{pmatrix} \overline{x} \\ \overline{p} \end{pmatrix}$$

$$K_1 = \begin{pmatrix} 0 & 1 \\ -k & 0 \end{pmatrix}$$

$$X_2 = \begin{pmatrix} \overline{x^2} \\ \overline{xp} \\ \overline{p^2} \end{pmatrix}$$

$$K_2 = \begin{pmatrix} 0 & 2 & 0 \\ -k & 0 & 1 \\ 0 & -2k & 0 \end{pmatrix}$$

$$X_3 = \begin{pmatrix} \overline{x^3} \\ \overline{x^2 p} \\ \overline{x p^2} \\ \overline{p^3} \end{pmatrix}$$

$$K_3 = \begin{pmatrix} 0 & 3 & 0 & 0 \\ -k & 0 & 2 & 0 \\ 0 & -2k & 0 & 1 \\ 0 & 0 & -3k & 0 \end{pmatrix}$$

$$X_4 = \begin{pmatrix} \overline{x^4} \\ \overline{x^3 p} \\ \overline{x^2 p^2} \\ \overline{x p^3} \\ \overline{p^4} \end{pmatrix}$$

$$K_4 = \begin{pmatrix} 0 & 4 & 0 & 0 & 0 \\ -k & 0 & 3 & 0 & 0 \\ 0 & -2k & 0 & 2 & 0 \\ 0 & 0 & -3k & 0 & 1 \\ 0 & 0 & 0 & -4k & 0 \end{pmatrix}$$

$$X_5 = \begin{pmatrix} \overline{x^5} \\ \overline{x^4 p} \\ \overline{x^3 p^2} \\ \overline{x^2 p^3} \\ \overline{x p^4} \\ \overline{p^5} \end{pmatrix}$$

$$K_5 = \begin{pmatrix} 0 & 5 & 0 & 0 & 0 & 0 \\ -k & 0 & 4 & 0 & 0 & 0 \\ 0 & -2k & 0 & 3 & 0 & 0 \\ 0 & 0 & -3k & 0 & 2 & 0 \\ 0 & 0 & 0 & -4k & 0 & 1 \\ 0 & 0 & 0 & 0 & -5k & 0 \end{pmatrix}$$

DIAGONAL FORMS OF MOMENT EQUATIONS

Either by direct computation or by decomposing K_n into operators similar to the creation and annihilation operators for Bosons and taking advantage of simple relationships between these operators, we get ($\omega^2 \equiv k$)

$$\left\{ \begin{array}{l} K_1^2 + \omega^2 = 0 \\ K_2(K_2^2 + 4\omega^2) = 0 \\ (K_3^2 + \omega^2)(K_3^2 + 9\omega^2) = 0 \\ K_4(K_4^2 + 4\omega^2)(K_4^2 + 16\omega^2) = 0 \\ (K_5^2 + \omega^2)(K_5^2 + 9\omega^2)(K_5^2 + 25\omega^2) = 0 \end{array} \right. \quad (8)$$

The regularities of these equations are obvious. These relations show that the moments satisfy the following non-matrix (diagonal) linear equations

$$\left\{ \begin{array}{l} X_1'' + \omega^2 X_1 = 0 \\ (X_2'' + 4\omega^2 X_2)' = X_2''' + 4\omega^2 X_2' = 0 \\ (X_3'' + 9\omega^2 X_3)'' + \omega^2 (X_3'' + 9\omega^2 X_3) \\ \quad = X_3'''' + 10\omega^2 X_3'' + 9\omega^4 X_3 = 0 \\ \left[(X_4'' + 16\omega^2 X_4)'' + 4\omega^2 (X_4'' + 16\omega^2 X_4) \right]' \\ \quad = X_4^{V} + 20\omega^2 X_4''' + 64\omega^4 X_4' = 0 \\ \left[(X_5'' + 25\omega^2 X_5)'' + 9\omega^2 (X_5'' + 25\omega^2 X_5) \right]'' \\ \quad + \omega^2 \left[(X_5'' + 25\omega^2 X_5)'' + 9\omega^2 (X_5'' + 25\omega^2 X_5) \right]' \\ \quad = X_5^{VI} + 35\omega^2 X_5'''' + 259\omega^4 X_5'' + 225\omega^6 X_5 = 0 \end{array} \right. \quad (9)$$

The equations for the lower moments are familiar. For example, X_1 equation gives

$$(\bar{x})'' + \omega^2 \bar{x} = 0 \quad (10)$$

which simply states that the center of gravity of the distribution oscillates as a single particle. The X_2 equation gives

$$(\bar{x}^2)'' + 4\omega^2 (\bar{x}^2)' = 0 \quad (11)$$

which is reminiscent of the equation for the Courant-Snyder β function.

BILINEAR INVARIANTS

We can define the bilinear invariants for the n th moment by

$$I_n \equiv \frac{1}{2} \tilde{X}_n S_n X_n \quad (12)$$

where \sim means transposition. The condition for invariance gives

$$\begin{aligned} 2 I_n' &= \tilde{X}_n' S_n X_n + \tilde{X}_n S_n X_n' \\ &= \tilde{X}_n \tilde{K}_n S_n X_n + \tilde{X}_n S_n K_n X_n = 0 \end{aligned} \quad (13)$$

or

$$\tilde{K}_n S_n + S_n K_n = 0 \quad (14)$$

It can be shown directly that S_n has the following forms

$$S_1 = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$$

$$\left\{ \begin{array}{l} S_2 = \begin{pmatrix} 0 & 0 & 1 \\ 0 & -2 & 0 \\ 1 & 0 & 0 \end{pmatrix} \\ S_3 = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & -3 & 0 \\ 0 & 3 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix} \\ S_4 = \begin{pmatrix} 0 & 0 & 0 & 0 & 1 \\ 0 & 0 & 0 & -4 & 0 \\ 0 & 0 & 6 & 0 & 0 \\ 0 & -4 & 0 & 0 & 0 \\ 1 & 0 & 0 & 0 & 0 \end{pmatrix} \\ S_5 = \begin{pmatrix} 0 & 0 & 0 & 0 & 0 & 1 \\ 0 & 0 & 0 & 0 & -5 & 0 \\ 0 & 0 & 0 & 10 & 0 & 0 \\ 0 & 0 & -10 & 0 & 0 & 0 \\ 0 & 5 & 0 & 0 & 0 & 0 \\ -1 & 0 & 0 & 0 & 0 & 0 \end{pmatrix} \end{array} \right. \quad (15)$$

The regularity is, again, obvious. They give the invariants

$$\left\{ \begin{array}{l} I_1 = 0 \\ I_2 = \overline{x^2} \overline{p^2} - (\overline{xp})^2 \\ I_3 = 0 \\ I_4 = \overline{x^4} \overline{p^4} - 4 \overline{x^3 p} \overline{xp^3} + 3 (\overline{x^2 p^2})^2 \\ I_5 = 0 \end{array} \right. \quad (16)$$

The invariant I_2 can be defined as a measure of the mean-squared emittance.

ENVELOPE EQUATIONS

We can define the n th envelope ξ_n (envelope of the n th moment) by

$$\xi_n \equiv \left(\overline{x^n} \right)^{\frac{1}{n}} \quad (17)$$

and derive second order equations for these envelopes.

(Strictly speaking, ξ_n can be interpreted as an "envelope" only for even values of n .) We shall demonstrate the procedure only for ξ_2

$$\begin{aligned} \left(\overline{x^2} \right)'' &= 2 \left(\overline{xp} \right)' = -2 \omega^2 \overline{x^2} + 2 \overline{p^2} \\ &= -2 \omega^2 \xi_2^2 + 2 \overline{p^2} \\ &= -2 \omega^2 \xi_2^2 + \frac{2}{\xi_2^2} \overline{x^2} \overline{p^2} \\ \left(\xi_2^2 \right)'' &= 2 \xi_2 \xi_2'' + 2 \left(\xi_2' \right)^2 \\ &= 2 \xi_2 \xi_2'' + \frac{1}{2 \xi_2^2} \left(2 \xi_2 \xi_2' \right)^2 \\ &= 2 \xi_2 \xi_2'' + \frac{1}{2 \xi_2^2} \left[\left(\overline{x^2} \right)' \right]^2 \\ &= 2 \xi_2 \xi_2'' + \frac{2}{\xi_2^2} \left(\overline{xp} \right)^2 \end{aligned}$$

Therefore

$$\xi_2'' + \omega^2 \xi_2 = \frac{\overline{x^2} \overline{p^2} - (\overline{xp})^2}{\xi_2^3} \equiv \frac{E_2}{\xi_2^3} \quad (18)$$

Similar procedure gives

$$\xi_n'' + \omega^2 \xi_n = (n-1) \frac{E_n}{\xi_n^{2n-1}} \quad (19)$$

where $\xi_n \equiv \left(\overline{x^n}\right)^{\frac{1}{n}}$ and

$$\begin{cases} E_1 = 0 \\ E_2 = \overline{x^2} \overline{p^2} - (\overline{xp})^2 \\ E_3 = \overline{x^3} \overline{xp^2} - (\overline{x^2 p})^2 \\ E_4 = \overline{x^4} \overline{x^2 p^2} - (\overline{x^3 p})^2 \\ E_5 = \overline{x^5} \overline{x^3 p^2} - (\overline{x^4 p})^2 \end{cases} \quad (20)$$

Only the equation for ξ_2 is useful, because E_2 is identical to the invariant I_2 . All other E_n obey rather complex time equations. The ξ_2 equation is reminiscent of the amplitude ($\omega \equiv \sqrt{\beta}$) equation of Courant and Snyder or the Kapchinsky-Vladimirsky equation in the absence of space charge. The ξ_2 equation may be called the rms envelope equation.

GENERALIZATIONS

Several straightforward generalizations should be pursued.

(a) When k is time dependent $k = k(t)$ we should write the single particle equation as

$$\begin{cases} x' = a(t)p, \\ p' = -b(t)x, \end{cases} \quad K_1 \equiv \begin{pmatrix} 0 & a \\ -b & 0 \end{pmatrix}, \quad \omega^2 \equiv ab \quad (21)$$

and proceed in a similar manner.

(b) When space-charge force is present we write

$$\begin{cases} x' = ap \\ p' = -bx + Fx \end{cases} \quad K_1 \equiv \begin{pmatrix} 0 & a \\ -b+F & 0 \end{pmatrix} \quad (22)$$

where $F = F(x,t)$ is the space-charge force. In this case a condition must be imposed on F to insure the invariance of I_2 . This generalization can proceed in the manner a la Sacherer.

(c) The generalization to more than one coupled dimensions can be made in a straightforward way as indicated by Sacherer.